



Impact of Centrifugal Modification of Magnetohydrodynamic Equilibrium on Resistive Wall Mode Stability

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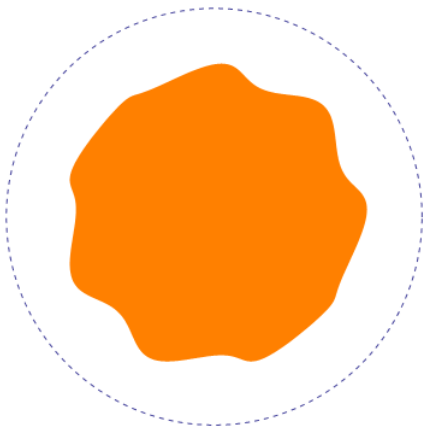
Summary

- ✓ RWMs in self-consistent (including rotational modification of the Grad-Shafranov equation) equilibria have been numerically investigated.
- ✓ Compared with conventional (using the static Grad-Shafranov equation) equilibria, in the self-consistent equilibria,
 - ✓ RWM growth rates are reduced for a wide parameter range of $\beta_N (=3\sim 5)$, rotation ($M^2 < \sim 0.15$) and wall location.
 - ✓ The stable window is enlarged and shifted.
 - ✓ Stable windows can exist even if the conventional equilibria have no window.
- ✓ The modification of equilibrium profile, not of eigenfunction, is essential to stabilization in self-consistent equilibria.
- ✓ In self-consistent equilibria, reduction of destabilizing energy δW_{uhp} and δW_{uhc} is essential to stabilization.

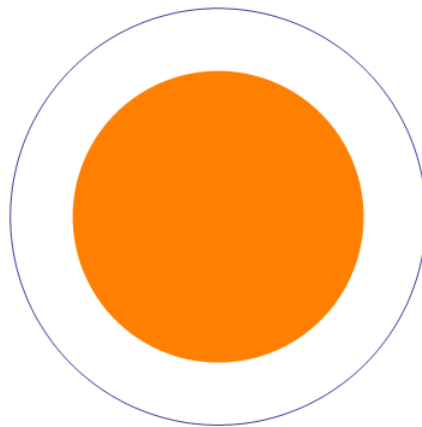
What is Resistive Wall Modes (RWMs)?

- ✓ RWMs originates from external kink modes ($\gamma\tau_A \sim 1$).
- ✓ Ideal walls stabilize external kink modes.
- ✓ Resistive walls slow down external kink modes to timescale of eddy current decay ($\gamma\tau_w \sim 1$).

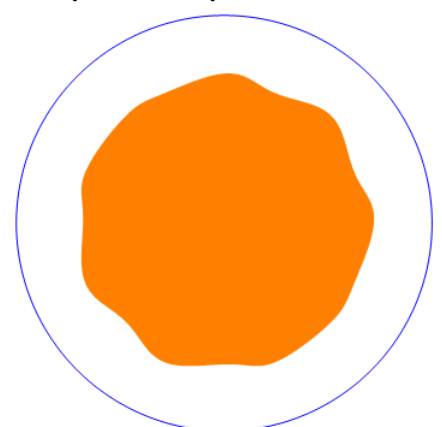
No wall (kink)



Ideal wall (stabilization)

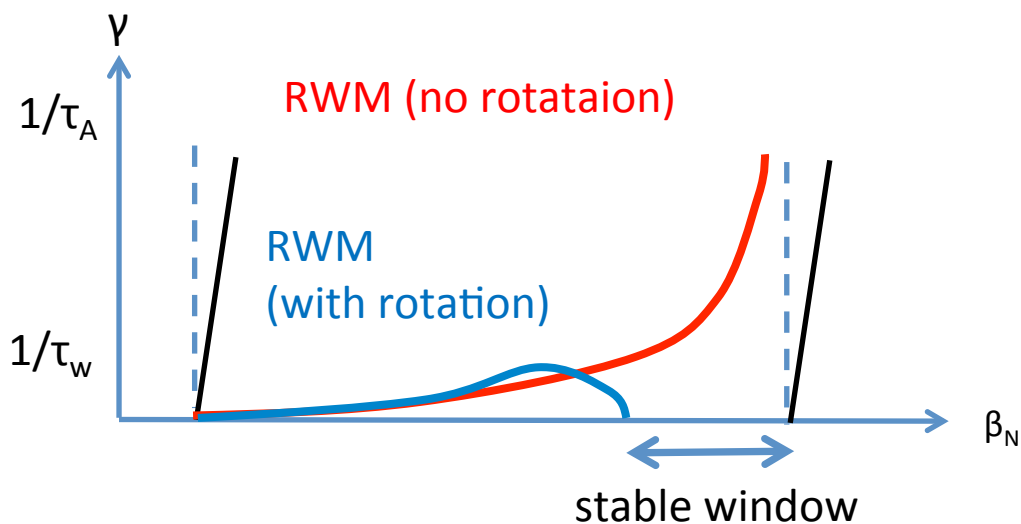


Resistive wall (RWMs)



Why RWM? How to stabilize RWM?

- ✓ Stabilization of RWMs is a necessary condition for operation of advanced tokamaks aiming at steady-state high- β_N plasma confinement such as JT-60SA.
- ✓ Many theoretical/experimental researches show rotational stabilization of RWMs.



As a basis of quantitative RWM study, we need to develop a numerical code for RWM in realistic tokamak geometry including plasma rotational effects.

RWM codes in tokamak geometry

- ✓ MARS-F (Chu PoP05), MARS-K (Liu NF09), CarMa (Liu PoP09)
 - ✓ Linearized resistive MHD, “perturbative” toroidal rotation, kinetic effects, 3D wall, feedback.
- ✓ NMA (Chu NF03)
 - ✓ Linearized ideal MHD without rotation, feedback.
- ✓ MISK (Berkery PRL11)
 - ✓ Linearized ideal MHD without rotation, kinetic effects
- ✓ VALEN (Bialek PoP01)
 - ✓ Linearized ideal MHD without rotation, 3D wall, and feedback
- ✓ **MINERVA(Aiba CPC09) with “RWMaC” modules**
 - ✓ We develop a new RWM code. It has some advantages :
 - ✓ (1) perturbative poloidal rotation (Aiba PoP11)
 - ✓ (2) centrifugal modification of MHD equilibrium by plasma toroidal rotation
 - ✓ (3) initial value approach

Rotational modification of MHD equilibrium

Under isothermal condition $T=T(\psi)$, existence of toroidal rotation generalizes the Grad-Shafranov equation as (e.g. Zehrfeld NF72)

$$\left\{ \begin{array}{l} \Delta^* \psi = -F \frac{dF}{d\psi} - \mu_0 R^2 \frac{\partial p}{\partial \psi} \Big|_R \\ p(\psi, R) = p_0(\psi) \exp \left[M^2(\psi) \left(\frac{R^2}{R_0^2} - 1 \right) \right] \end{array} \right. \quad (1)$$

Definition of two types of equilibria used in this poster

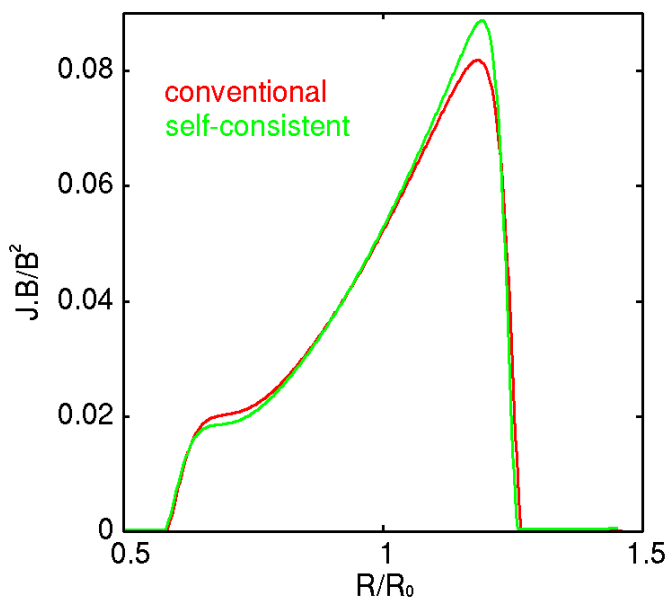
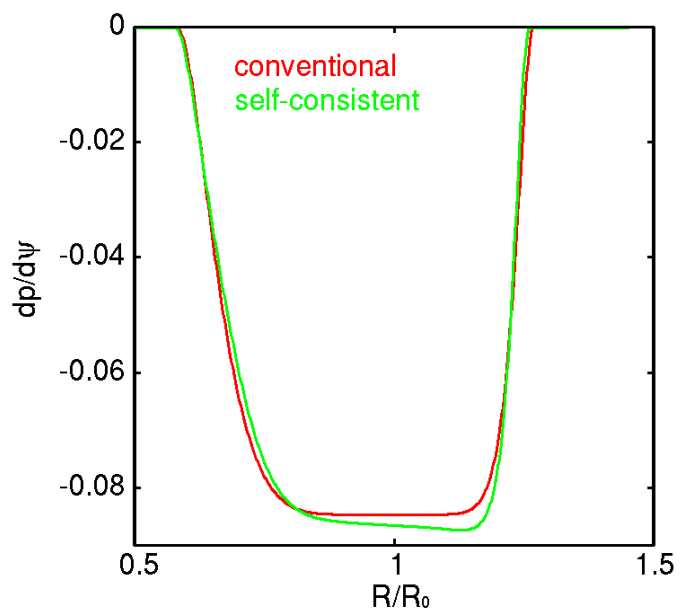
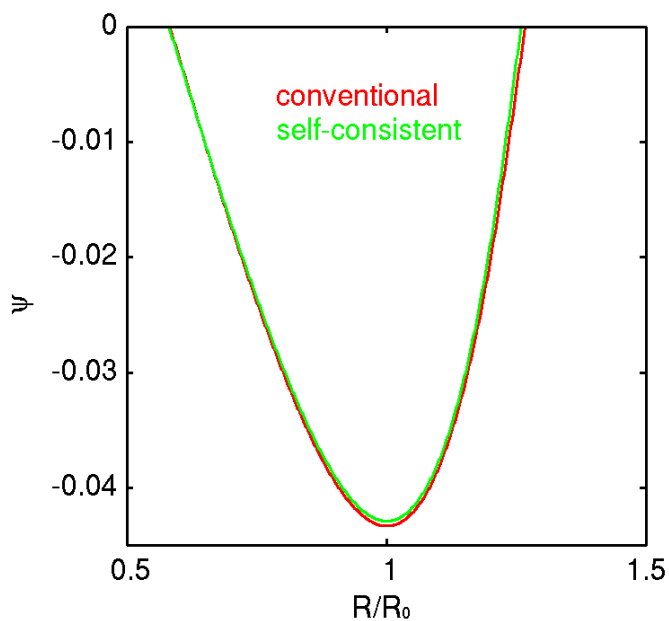
We call the solution to (1) as “**self-consistent equilibrium.**”

Conventional studies approximate $p \sim p_0$ due to the smallness of M^2 , i.e., the equilibrium is static.

We call this approximated solution as “**conventional equilibrium.**”

Rotational modification of MHD equilibrium (cont.)

Examples for “conventional” and “self-consistent” equilibria with $\beta_N \sim 2.83$ and $M^2 = 0.1$



Note : Pressure gradient and parallel current are affected by self-consistent inclusion of rotation compared with ψ .

Formalism for linear dynamics of RWMs – vacuum and resistive wall dynamics

Ampère's law across the resistive wall

$$\chi^{(w+)}(\theta, \phi, t) - \chi^{(w-)}(\theta, \phi, t) = \mu_0 \kappa \quad (2)$$

Faraday's and Ohm's laws across the resistive wall

$$\frac{\Delta |\nabla_S|}{\eta} \frac{\partial \tilde{B}^{(n)}}{\partial t} = -\nabla \cdot \left(|\nabla_S|^2 \nabla_{\perp} \kappa \right) \quad (3)$$

Quadratic form from (2) and (3) yields energy balance

$$\delta W_{IV} + \delta W_{OV} + \frac{1}{2\mu_0} \int_{S_p} \chi^{(p+)} \mathbf{Q}_e^* \cdot \hat{n} dS + \delta D_w = 0 \quad (4)$$

plasma response

vacuum magnetic energy $\delta W_{IV(OV)} = \frac{1}{2\mu_0} \int_{IV(OV)} |\nabla \chi^{\pm}|^2 d\tau$

energy dissipation in resistive wall $\delta D_w = \frac{1}{2} \int_{\text{wall}} \kappa \mathbf{B} \cdot d\mathbf{S}$

Formalism for linear dynamics of RWMs – linear plasma response

Plasma is governed by the Frieman-Rosenbluth equation (Frieman RMP60), which is the linearized ideal MHD with equilibrium rotation.

$$\rho \partial_t^2 \xi + 2\rho(\mathbf{u} \cdot \nabla) \partial_t \xi = (F_s + F_d) \xi \quad (5)$$

static and dynamic force operator

Linear plasma response of the Frieman-Rosenbluth equation reads

$$\frac{1}{2\mu_0} \int_{S_p} \chi^{(p+)} \mathbf{Q}_e^* \cdot \hat{n} dS = \delta K + 2\delta W_c + \delta W_p \quad (6)$$

Kinetic energy : $\delta K = \frac{1}{2} \int \xi^* \cdot \rho \partial_t^2 \xi d\tau$

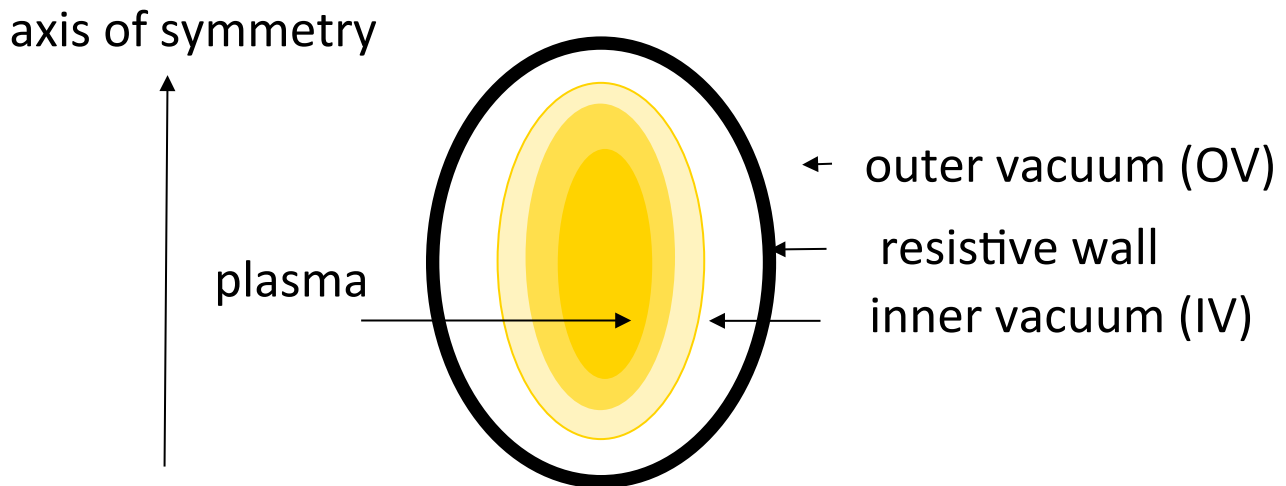
convective energy : $\delta W_c = \frac{1}{2} \int \xi^* \cdot \rho(\mathbf{u} \cdot \nabla) \partial_t \xi d\tau$

Potential energy including equilibrium rotation

$$\delta W_p = \frac{1}{2} \int \xi^* \cdot \left[\frac{|\mathbf{Q}|^2}{\mu_0} + \mathbf{J} \cdot (\xi^* \times \mathbf{Q}) + (\xi \cdot \nabla p) \nabla \cdot \xi^* + \Gamma p |\nabla \cdot \xi|^2 - \rho \xi (\mathbf{u} \cdot \nabla) \mathbf{u} + \rho \mathbf{u} (\mathbf{u} \cdot \nabla) \xi \right] d\tau \quad (7)$$

Note : “Conventional equilibrium” approach introduces equilibrium rotation \mathbf{u} in (6). The equilibrium quantities such as \mathbf{B} , p , and ρ include the modification induced by toroidal rotation, which is based on the solution of “static” equilibrium.

Formalism for linear dynamics of RWMs – energy balance



(4) and (6) yield energy balance in plasma-wall-vacuum system.

$$\begin{array}{c} \text{plasma} \\ \delta K + 2\delta W_c + \delta W_p \end{array} + \begin{array}{c} \text{vacuum} \\ \delta W_{OV} + \delta W_{IV} \end{array} + \begin{array}{c} \text{wall} \\ \delta D_W \end{array} = 0 \quad (8)$$

kinetic energy convective energy potential energy with rotation vacuum magnetic energy in IV/OV energy dissipation in resistive wall

RWM dynamics is governed by the balance among these energy sources (sinks).

“RWMaC” modules compute $\delta W_{IV(OV)}$ and δD_w

To solve (8) we have developed “RWMaC” modules to compute $\delta W_{IV(OV)}$ and δD_w , and have implemented them into MINERVA (Aiba CCP09) that computes δK , δW_p , and δW_c .

Inner vacuum and outer vacuum : $\delta W_{IV(OV)}$

Governing equation : Laplace equation for χ
(magnetic scalar potential $\mathbf{B} = \nabla \chi$)

Numerical scheme : hybrid FEM for IV
hybrid FEM or Green’s function
method for OV

Resistive wall : δD_w

Governing equation : diffusion equation for κ
[current potential $\mathbf{J} = (\nabla s \times \nabla \kappa) \delta(s - s_w)$]

Numerical scheme : hybrid FEM

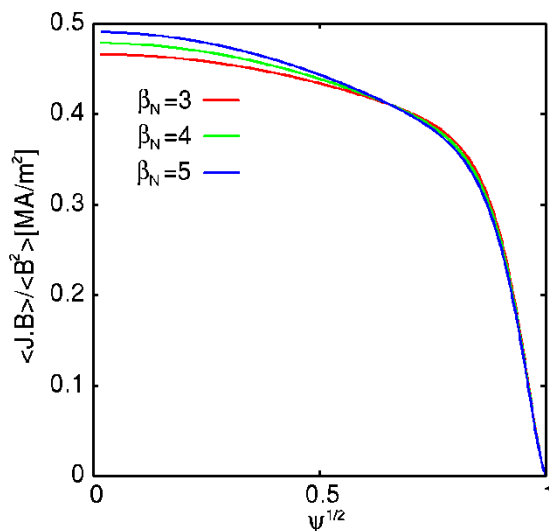
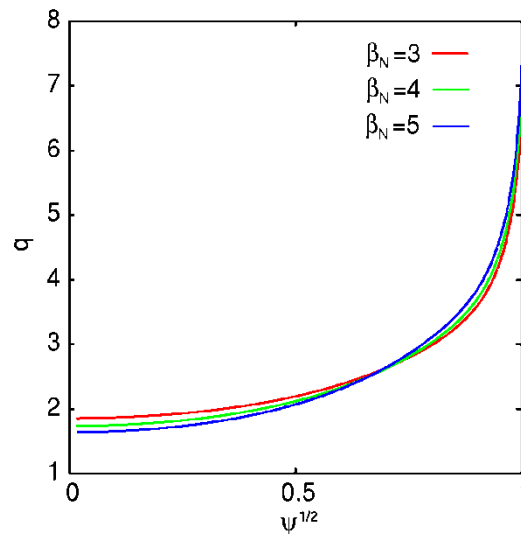
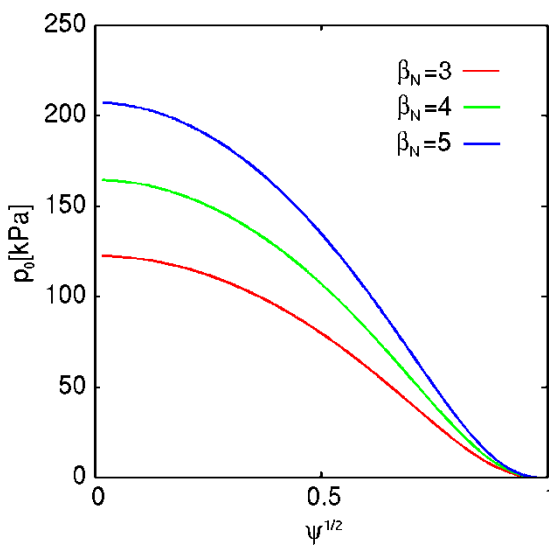
Boundary conditions on resistive wall and plasma surface :

Continuity of normal magnetic field + natural boundary
condition

RWMs in self-consistent equilibria – high- β_N equilibria for JT-60SA

By MINERVA/RWMAc, we can study the RWMs in self-consistent equilibria.

We consider high- β_N ($2.8 < \beta_N < 5.5$) equilibria with fixing D-shape of plasma surface ($\kappa=1.91$ and $\delta = 0.5$), toroidal magnetic field $B_0=1.7T$, and plasma current $I_p=2.3MA$, which are typical parameters for advanced plasma designed for JT-60SA.

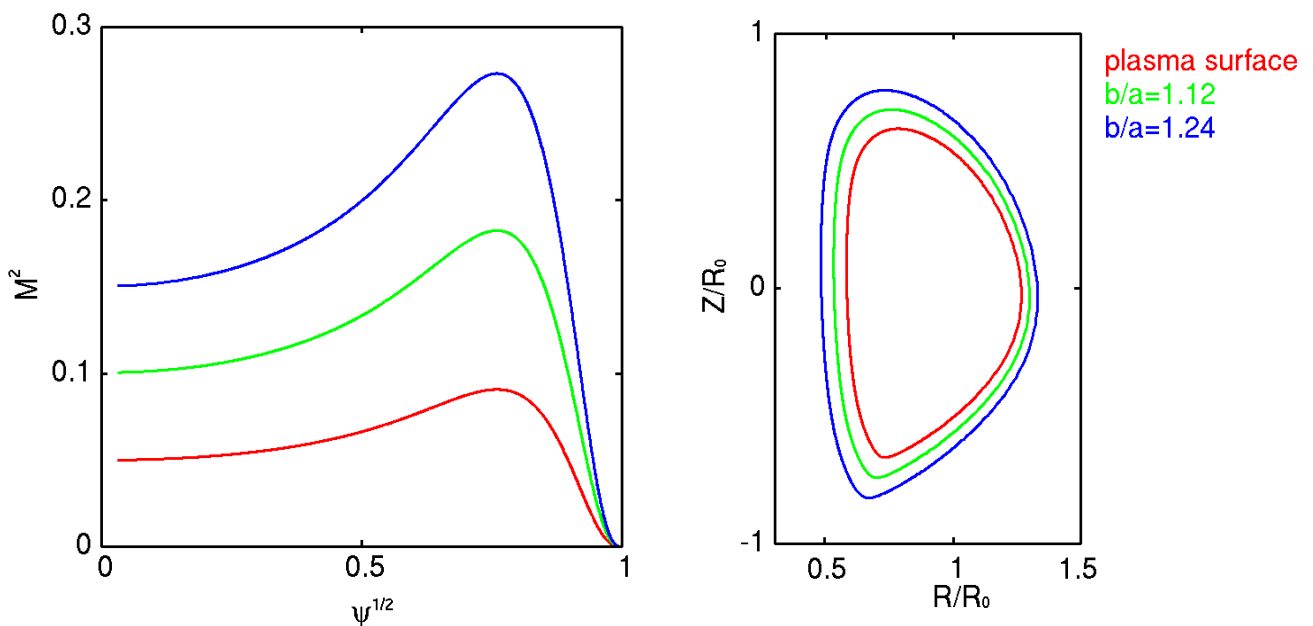


β_N is increased by scaling p_0 with keeping the almost same profiles of safety factor and parallel current.

RWMs in self-consistent equilibria – wall location rotation profiles

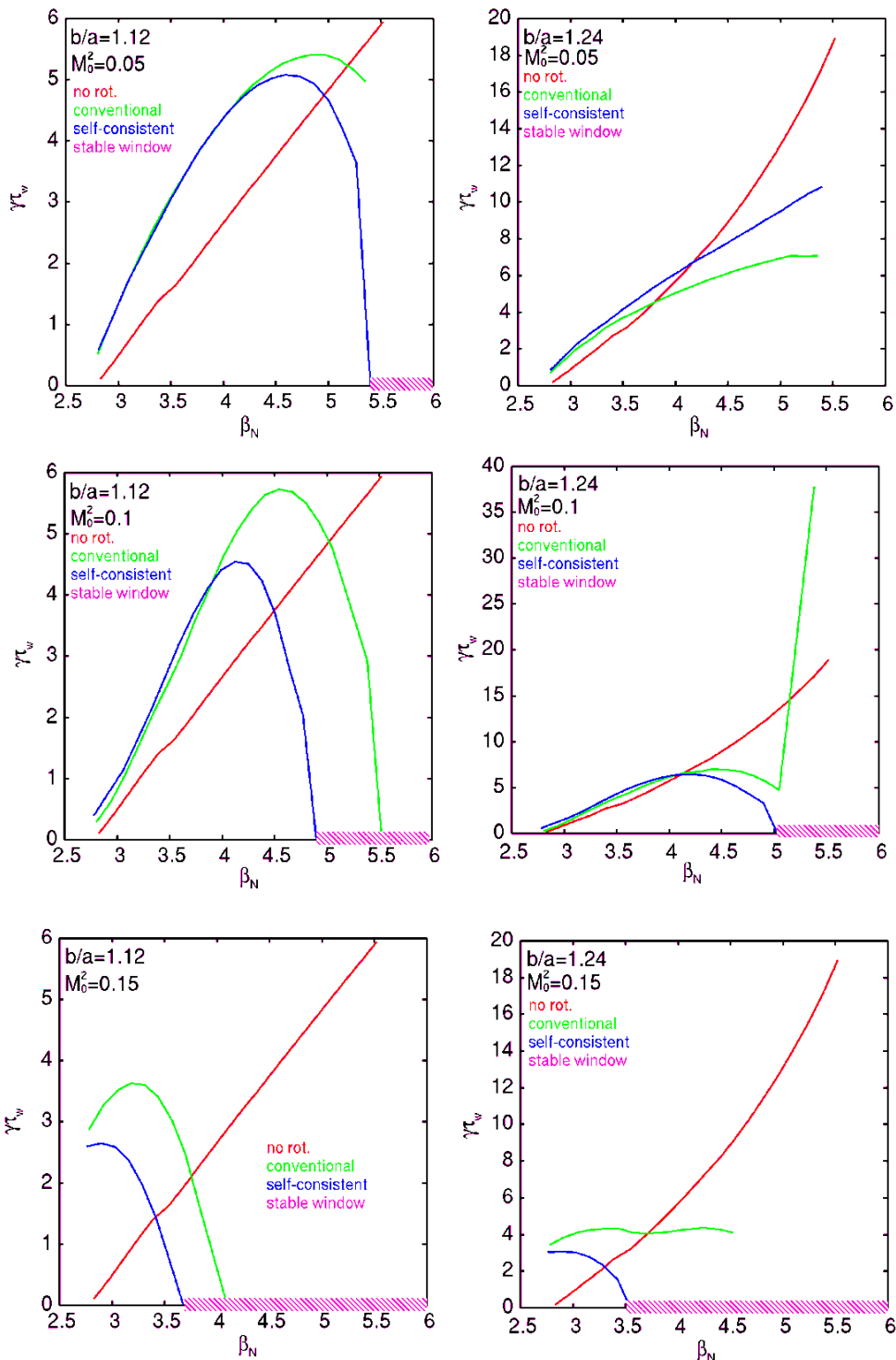
Rotation profiles is characterized by $\Omega(\psi)=\Omega_0[1-(\psi/\psi_{\text{sur}})^5]^2$. By changing Ω_0 and fixing temperature at magnetic axis, we consider 3 cases of squared Mach number rotation as $M^2=0.05, 0.1$, and 0.15 which are relevant to low-aspect ratio tokamaks.

We use two wall locations $b/a=1.12$ and 1.24 .



RWM stabilization in self-consistent equilibria – scan by β_N , rotation, and wall location

Normalized RWM growth rates without rotation, in **conventional** equilibrium, and **self-consistent** equilibrium as functions of β_N



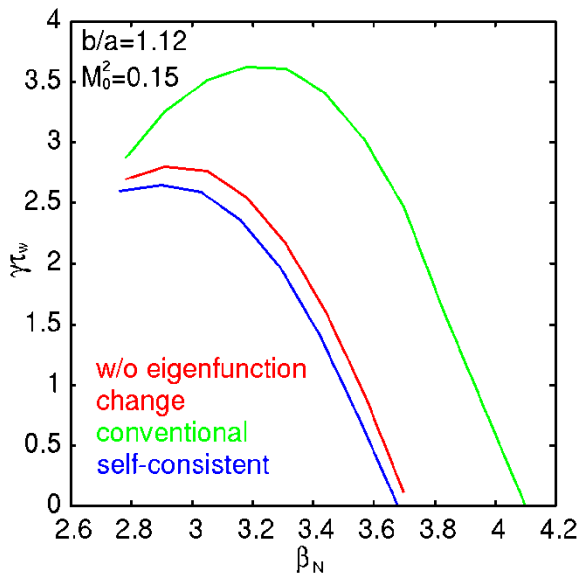
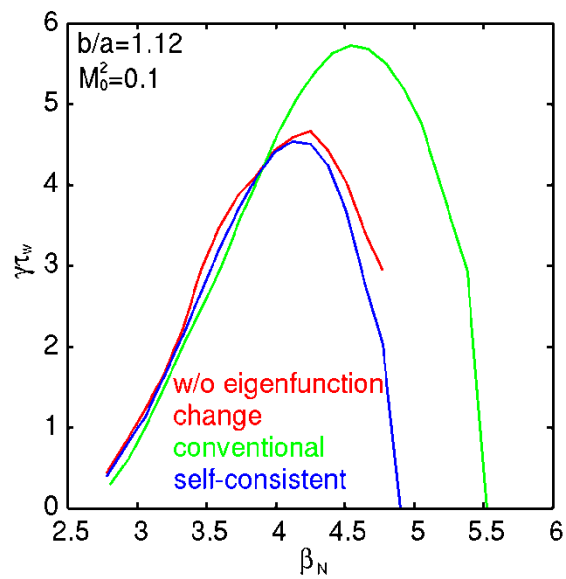
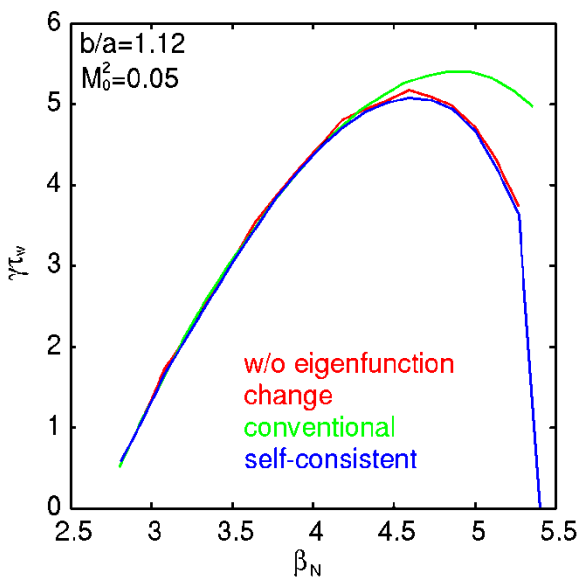
Note on scan by β_N , rotation, and wall location

- ✓ RWM growth rates are reduced in a wide range of wall location and β_N .
- ✓ Some cases ($b/a=1.12$ and $M^2=0.1, 0.15$) show that
 - ✓ the self-consistent equilibrium has an extended stable window.
 - ✓ the location of stable window is shifted.
- ✓ Other cases ($b/a=1.12$ and $M^2=0.05$, $b/a=1.24$ and $M^2=0.1, 0.15$) show that
 - ✓ The self-consistent equilibrium has a stable window even if the conventional equilibrium has no window.

In self-consistent equilibria, RWMs are stabilized by equilibrium change self-consistently introduced in the equilibrium.

Modification of equilibrium is essential to RWM stability – eigenfunction modification is not

Rotation modifies eigenfunction as well as equilibrium.



RWM problem reads $A\xi = \lambda B\xi$.
 Defining Δ by $\Delta f = f_s - f_c$ where s (c) indicates self-consistent (conventional), the problem reads

$$(A_c - \lambda_c B_c) \Delta \xi + B_c \Delta (B^{-1} A) \xi_s = \Delta \lambda B_c \xi_s$$

| | | |
|----------------------|--------------------|-------------------|
| eigenfunction change | equilibrium change | eigenvalue change |
|----------------------|--------------------|-------------------|

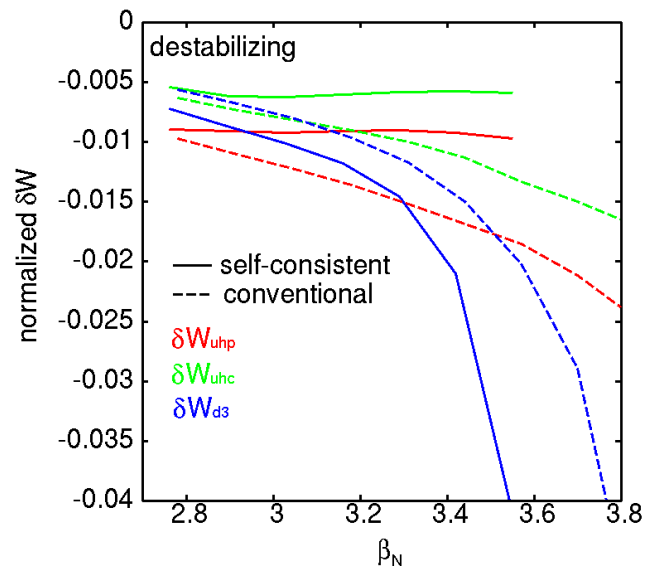
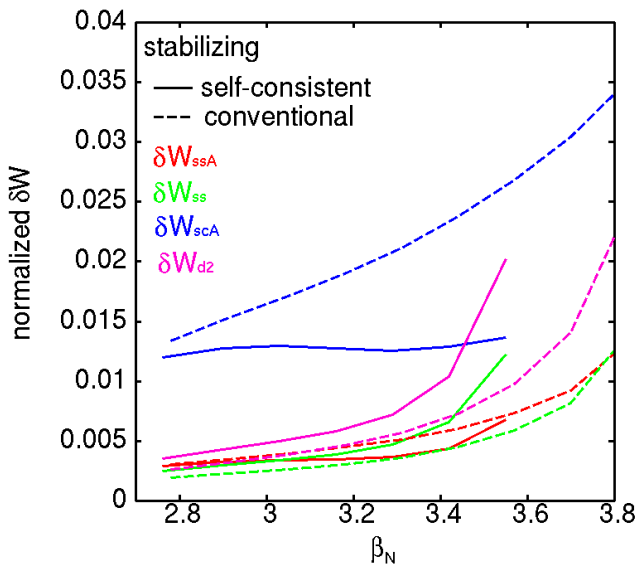
Reduction of pressure and current driven terms is essential to RWM stabilization

From energy balance (8), we get

$$\gamma = \text{Re}(\lambda) = - \frac{\delta W_p + \delta W_{IV} + \delta W_{OV}}{|\delta D_w + 2\delta W_c|^2} \delta D_w$$

We decompose potential energy δW_p as

$$\begin{aligned} \delta W_p &= \delta W_{ssA} + \delta W_{ss} + \delta W_{scA} && \text{: stabilizing} \\ &+ \delta W_{uhp} + \delta W_{uhc} && \text{: destabilizing by pressure and current driven} \\ &+ \delta W_{d1} + \delta W_{d2} + \delta W_{d3} && \text{: mode compression, shear, centrifugal effect due to rotation} \end{aligned}$$



In self-consistent equilibrium, we get smaller δW_{scA} and larger δW_{d3} . So it seems to lead to destabilization, however, reduction of destabilizing δW_{uhp} and δW_{uhc} leads to stabilization